

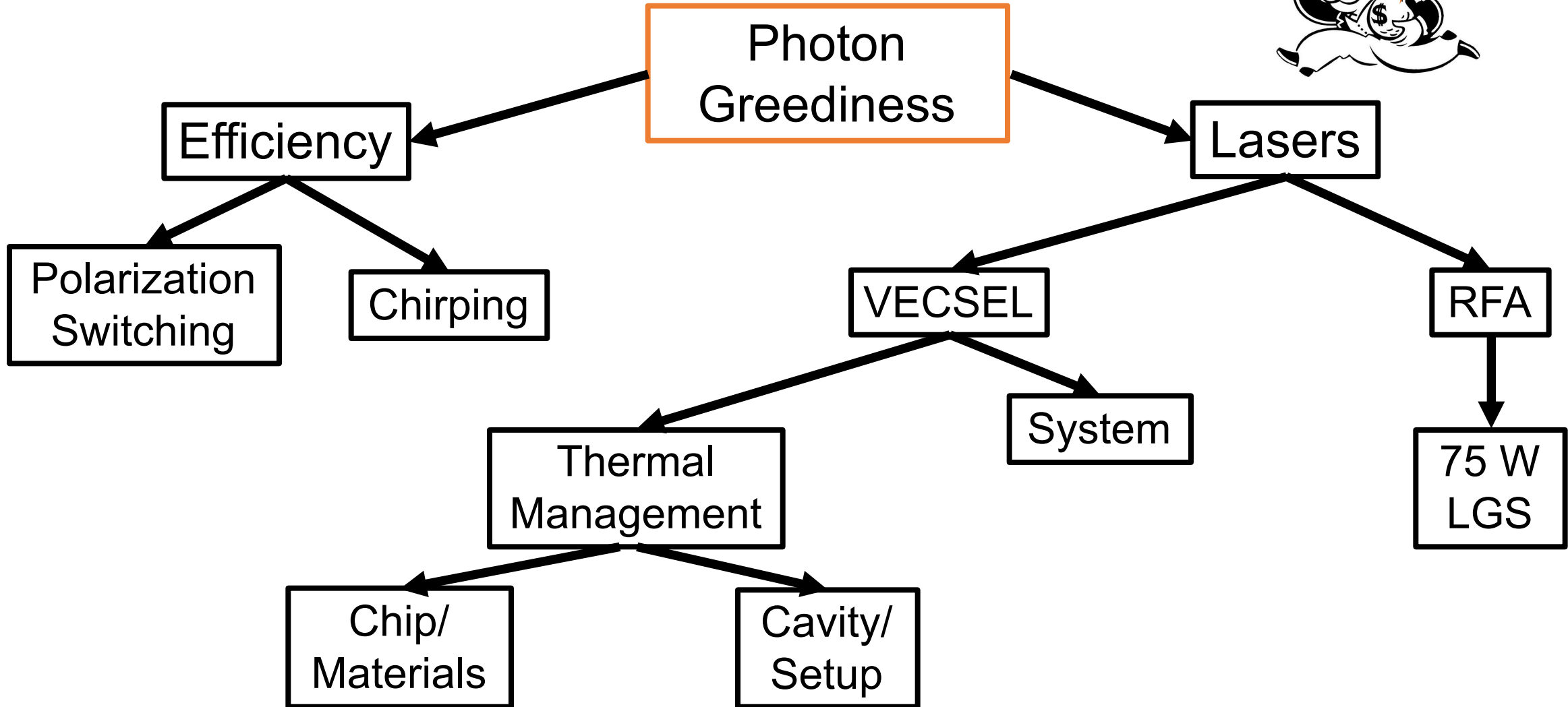


# Sodium Beacon VECSELS and SOR Laser Overview

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AFRL/RDSS 11 June 2019

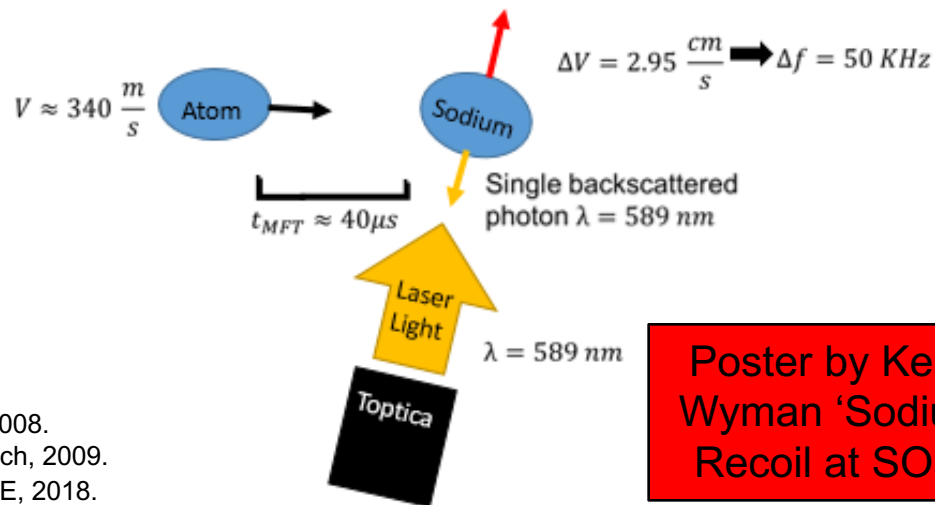
# Next Generation at SOR



# Efficiency

## Chirping

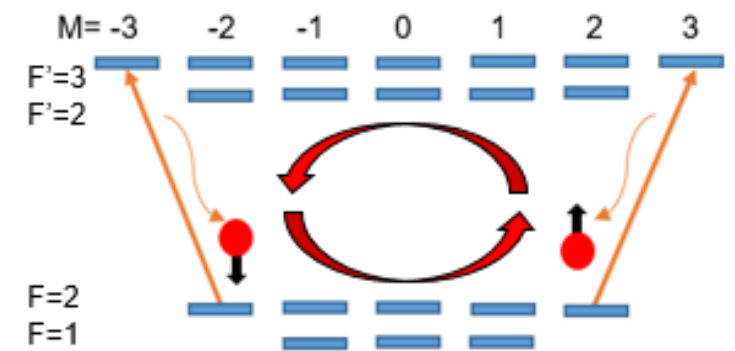
- Recoil/Radiation Pressure
- Red-shifted 50 kHz based on Doppler effect
- Stops with: Collision, spin exchange or inability of source radiation to keep up
- Rate depends on radiative lifetime and fraction of atoms in  $^2P$  state
- Chirping over the length of a mean free time could increase photon returns



Poster by Keith Wyman 'Sodium Recoil at SOR'

## Polarization Switching

- Larmor precession due to the Earth's geomagnetic field (0.48G)
- Redistributes the magnetic sub-level population
- Negates  $F = 2$  to  $F = 3$  transition of optical pumping if angle between the laser and the geomagnetic field is large enough
- Switch beam polarization at the Larmor frequency ( $\sim 328\text{kHz}$ ) to trap atom between two ideal repumping states



Moussaoui, N., et al, Astron. Astrophys., 501, pp. 793799, 2009.  
 Milonni, P. W., et al, J. Opt. Soc. Am., A 16, pp. 25552566, 1999.  
 Drummond, J., AMOS Conference Proceedings, September 2007.  
 Holzlöhner, R., et al, A and A 510, A20, October 2009.

# Efficiency

## Chirping

- Recoil/Radiation Pressure
- Red-shift
- Stops with
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- Rate dep
- fraction of
- Chirping
- could inc

$$v \approx 340 \frac{m}{s}$$

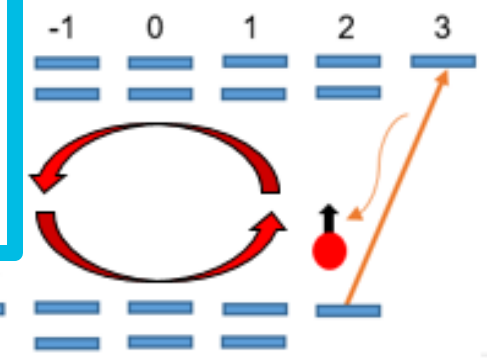
## Polarization Switching

- Larmor precession due to the Earth's

### LIDAR

- 24-7 atmospheric monitoring
- Includes detectors, filters, data acquisition managers, optics, and pulsed 3-frequency solid state laser system
- Center wavelength 589nm
- Un-conditioned environments w/ minimal human intervention
- Signal to noise ratio greater than 5 in day and night conditions
- Reasonable cost for small observatories
- Monitoring of sodium layer height, wind, temperature

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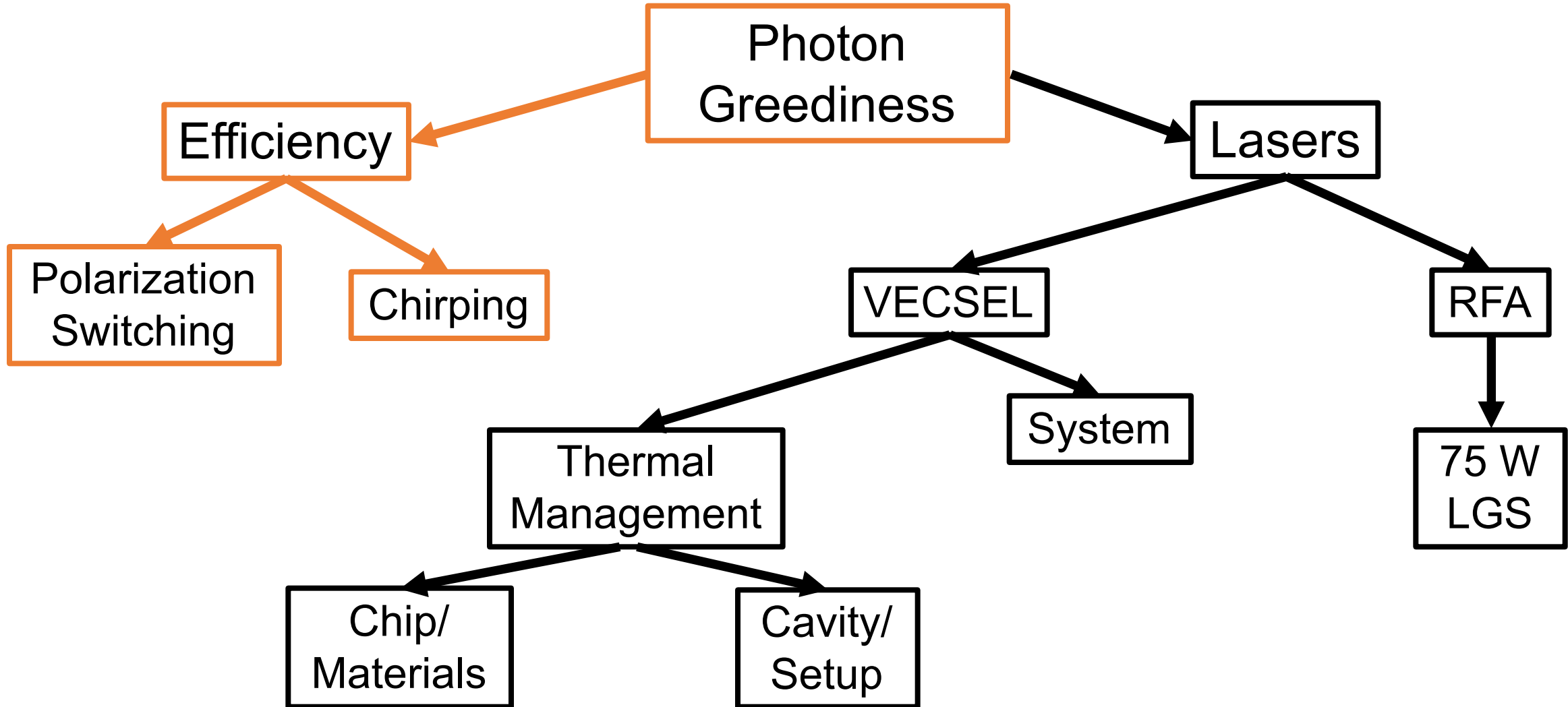


Poster by Keith Wyman 'Sodium Recoil at SOR'

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Hillman, et al, SPIE, 2008.  
 Kibblewhite, AMOSTech, 2009.  
 Wyman, K., et al, SPIE, 2018.  
 Hackett, Shawn, UNM Digital Repository, November 2016

# Next Generation at SOR



# 75 W Laser Beacon

## Two-Phase OTA :

1. RFA risk reduction- 15 months
2. 75 W beacon prototype- 27 months

## Technical goals for Phase II:

- Laser power of at least 75 W at 589.159 nm
- Re-pumper 1717.8±10 MHz away from laser line center with at least 22.5% power
- Tunable across entire D2a line
- Bandwidth 12.5 MHz FWHM or less
- Wavelength stabilized to peak of sodium line by ±10 MHz or better
- Include frequency chirping

Date closes: 21 June 2019

Located on FedBizOpps, Solicitation Number: **FA9451-19-9-0001**



# VECSEL effort overview

VECSEL, OPSSL, SDL, etc  
 different from VCSEL  
 diffraction-limited beam and higher optical powers

## Benefits of VECSELS

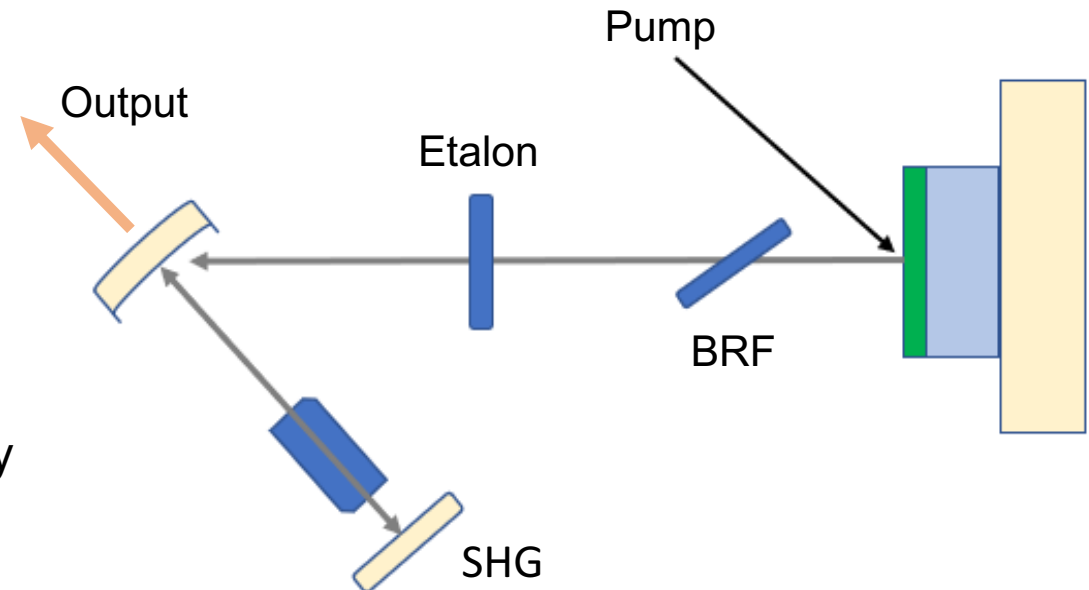
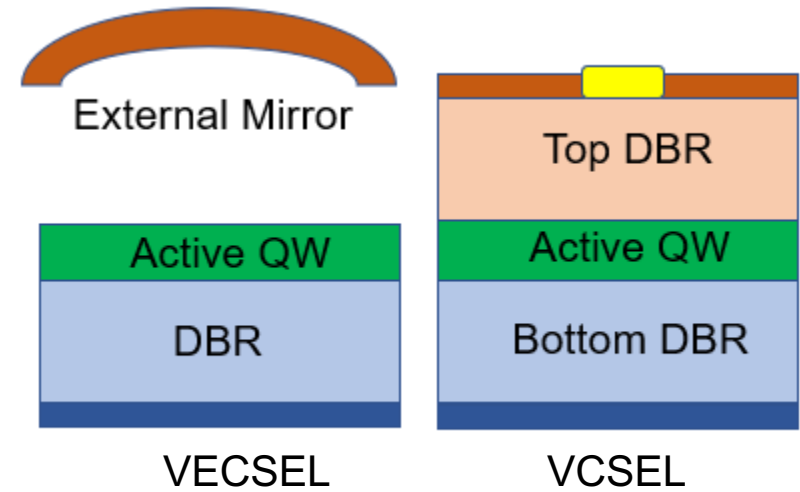
low cost  
 low complexity and compact  
 reliable, little maintenance

## Potential use as next-generation beacon

beam combination, multiple beacons, or small observatories  
 dedicated repumper

## Difficulties with VECSELS

difficult to get high power and single-frequency  
 thermal management issues



# VECSEL effort overview

## STTR Overview

Current Phase II of a Small Business Technology Transfer (STTR) program to develop Vertical External-Cavity Surface-Emitting Laser (VECSEL) technology

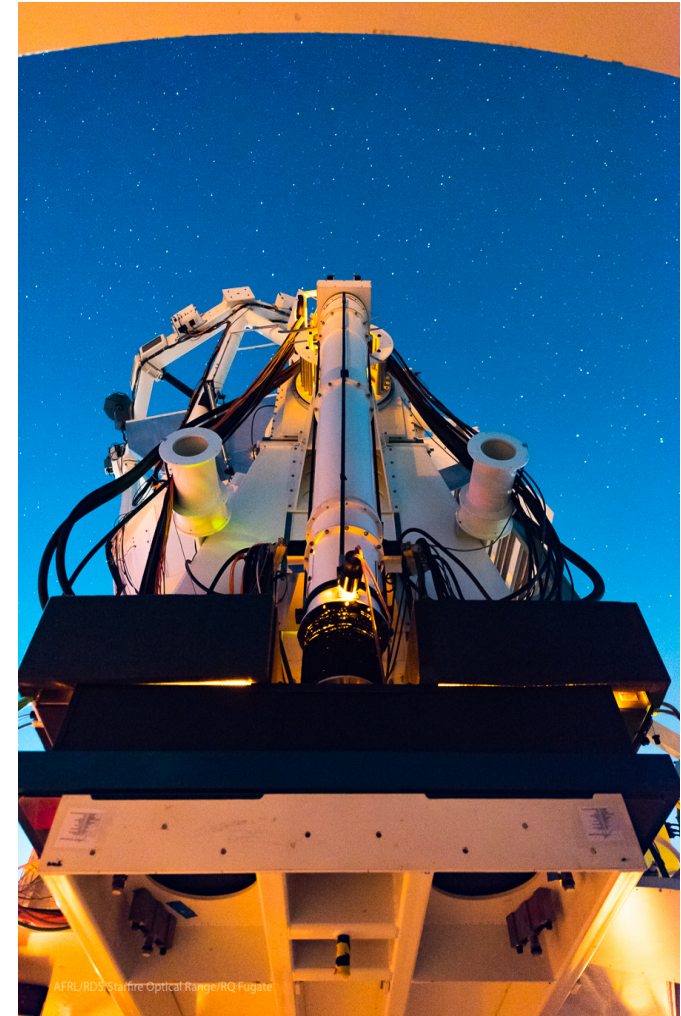
- The STTR program helps provide financial assistance to pursue technologies in support of DoD missions
- Encourages collaboration between industry and research institution

## This STTR

For a polychromatic VECSEL at 589nm and 1140nm

- Phase I: Develop plan for <1GHz, 6GHz tunable, >10W VECSELS at 1140 and 589nm
- Phase II: Develop lab demonstration with  $\lambda$  stability, system robustness, and proof of concept implementation on-sky
- Phase III: Deliver working prototype capable of attaching to telescope for use at 1140 and 589nm

5 or 6 months into the effort

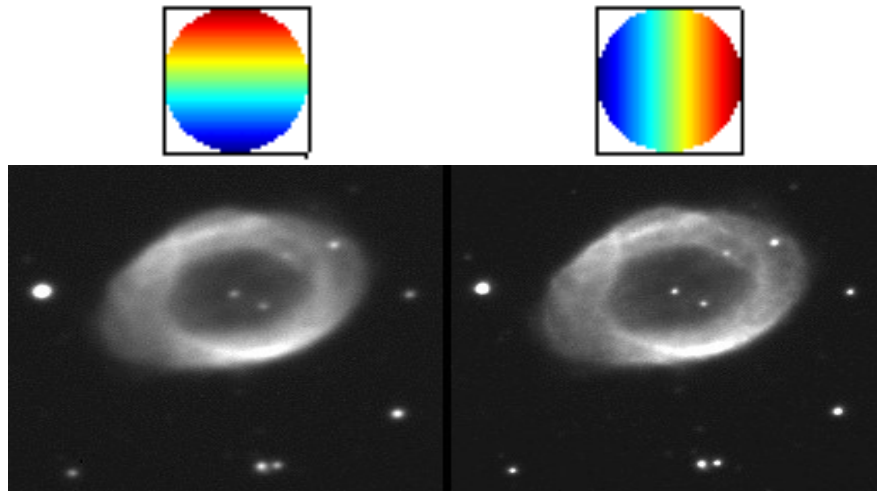




# Polychromatic laser guide stars

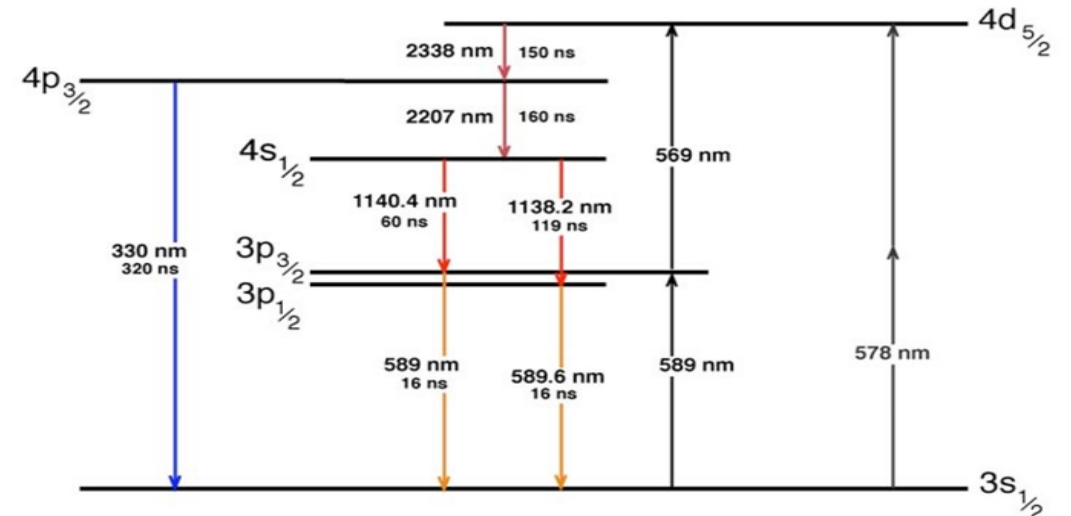
- Tip-tilt aberration, cannot be corrected by monochromatic sources, same optical path
- Typically due to atmospheric refraction, causes image smearing for most relevant objects (planets, satellites, etc)
- Usually a separate two-axes, tip-tilt mirror looking at a NGS
- Polychromatic guide star can be used (overlapping, different optical path)

1140 nm is chosen because atmospheric transmission, difference in index of refraction, source availability, limits Rayleigh need >10W in both wavelengths

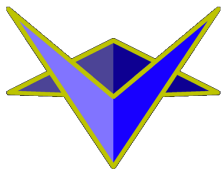


Imagery of the Ring Nebula taken with and without tip-tilt AO correction  
*S.B.I.G. Adaptive Optics System Introduction; September 1997*

Pumps the sodium  $3P_{3/2}$  to  $4S_{1/2}$  transition

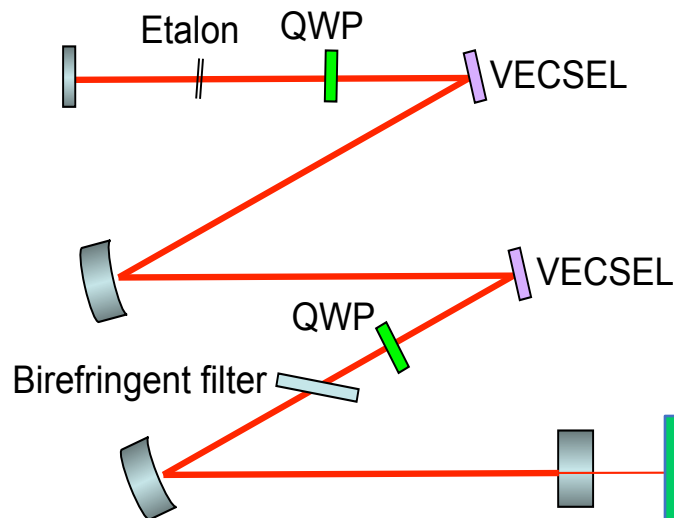
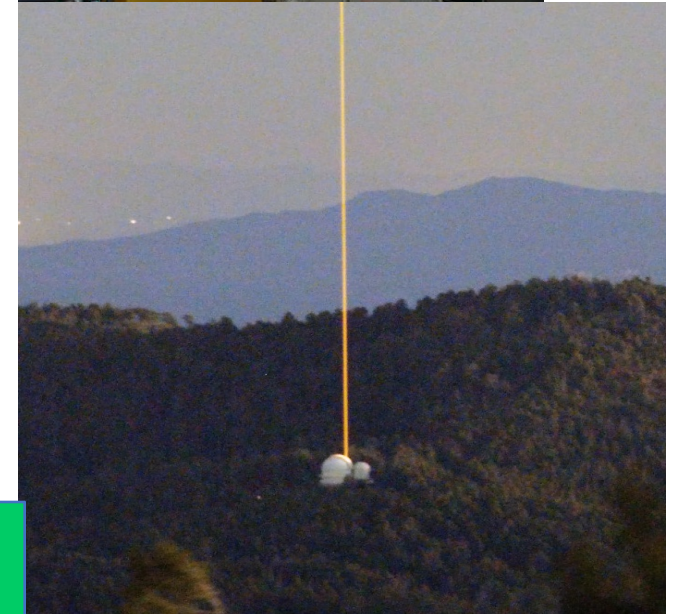
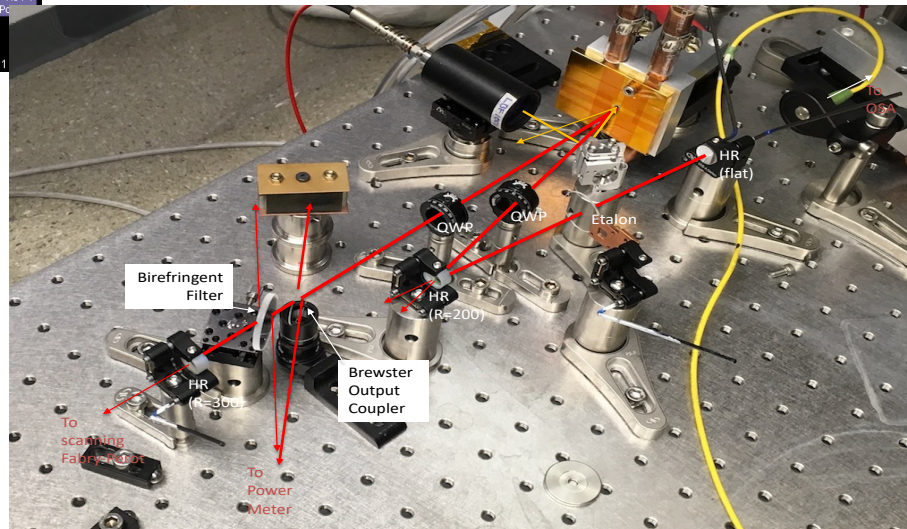
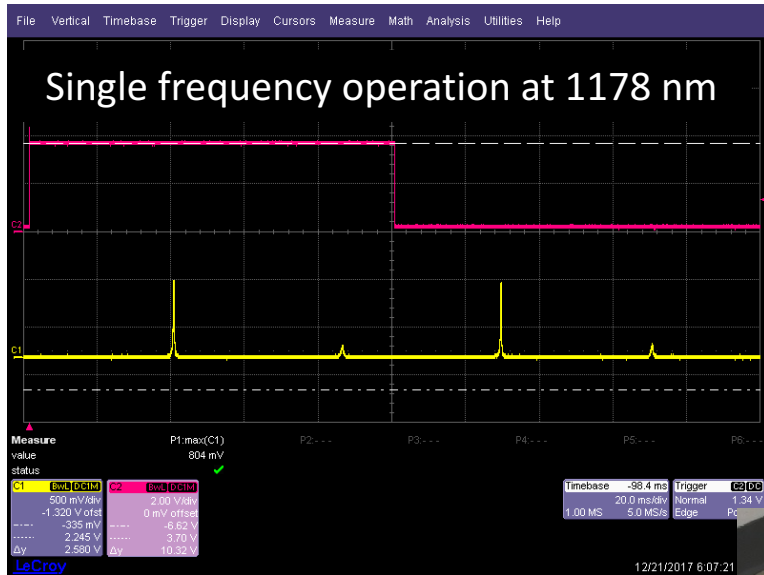


Hackett, Shawn, UNM Digital Repository, November 2016



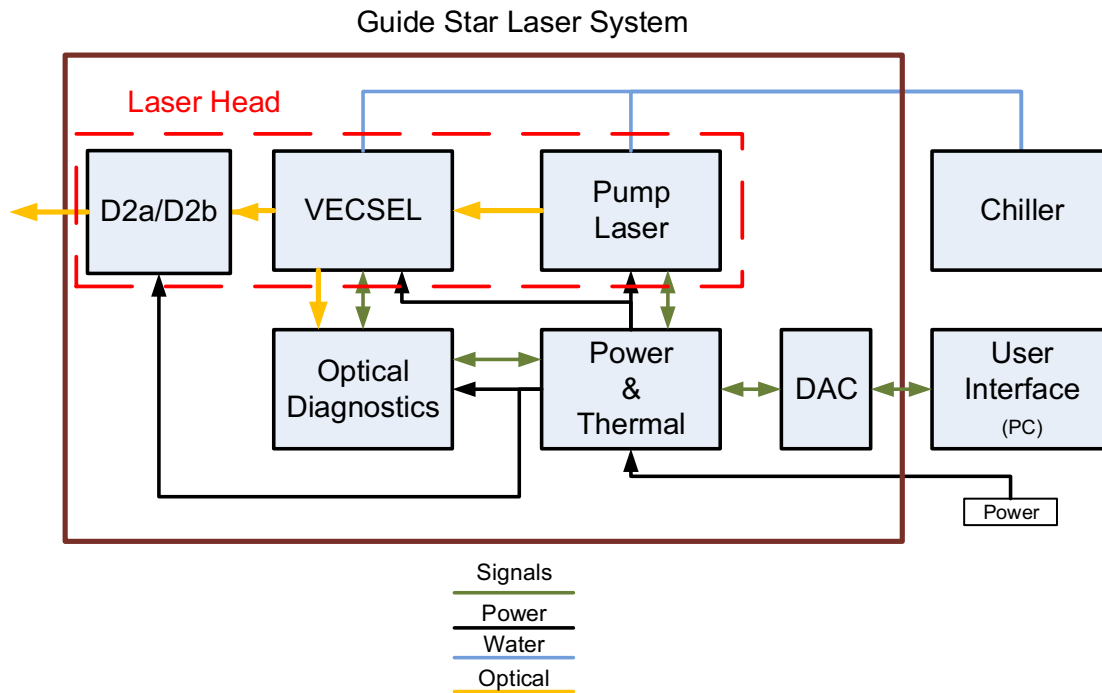
# Single-frequency VECSEL

- The HartSCI/UA approach uses a “twisted mode” cavity – orthogonal circular polarizations in forward and reverse directions.
- Enforces single-line frequency stability with multiple VECSEL devices in the cavity for longitudinal power scaling.
- Currently running at 1178 nm; will use SHG to get to 589 nm.
- Phase II will see a 4-VECSEL laser with SHG fielded at UA’s 1.5 m Kuiper telescope.



**Achieved > 10 W single frequency with 2-VECSEL laser at 1178 nm**

# Areté's VECSEL GSL System

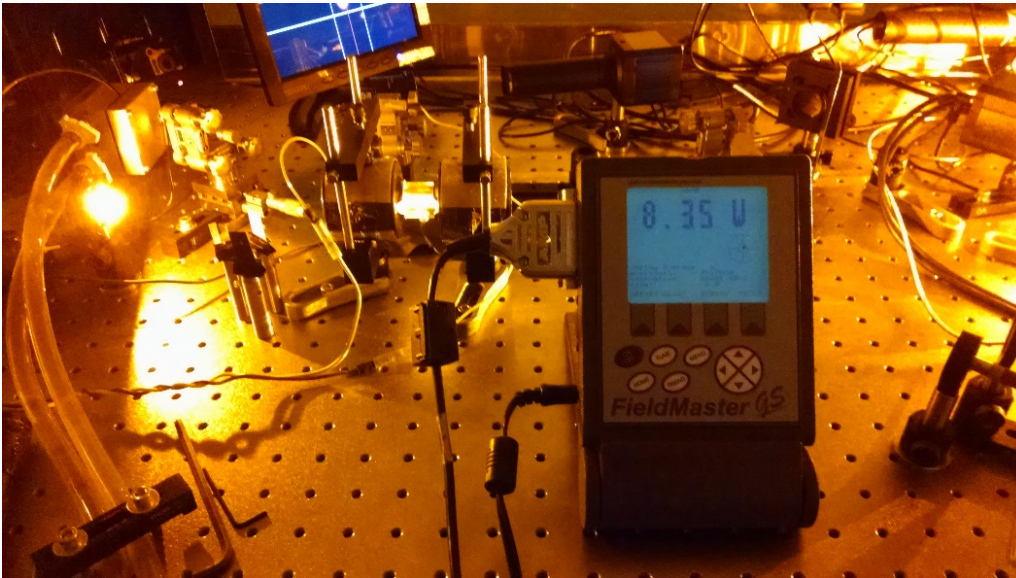


Goal: Design a system that:

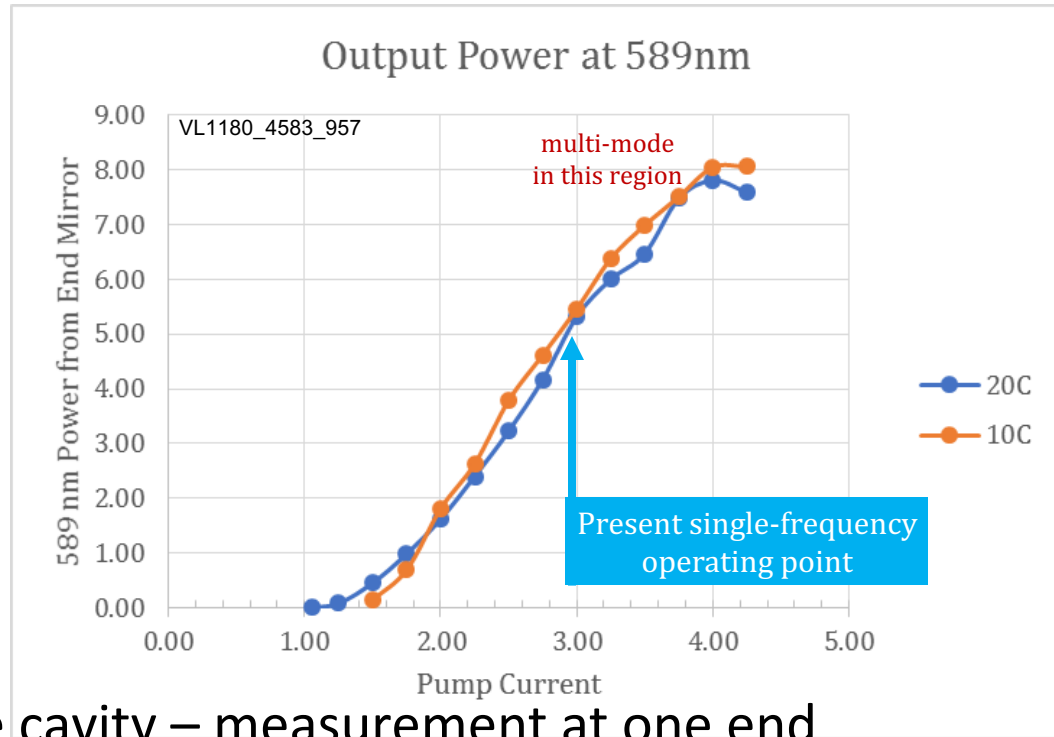
- Demonstrates the viability of VECSELs for Guidestar applications
- Serves as a foundational prototype on which to build future units
  - Lowers acquisition and maintenance costs of GSLs
- Provides utility to astronomy, space situational awareness, communications, and other applications

Characteristic	Values and Rationale
Primary $\lambda$ and Power	8-20 W locked to Na(D <sub>2a</sub> ) ~589 nm
Secondary $\lambda$ and Power	Tunable and lockable at D <sub>2b</sub> $\Delta\lambda = 1.7$ GHz from D <sub>2a</sub>
Waveform	Continuous Wave
Linewidth	5-50 MHz
Fine Tuning	~1 GHz, continuous <i>Scan sodium transition to enable line locking</i>
Gross Tuning	~5 GHz, does not need to be continuous <i>Allow capture of Rayleigh backscatter</i>
Beam Quality	$M^2 < 1.2$ <i>Near Diffraction Limited</i>
Polarization	Well defined polarization, contrast ratio >20 <i>Circular polarization is broadcast</i>
User Interface	PC Based GUI
Diagnostics	Wavelength and power
Power	110-240 V AC
Water	4-8 slpm flow of <i>cool</i> water

## Power from End Mirror Only

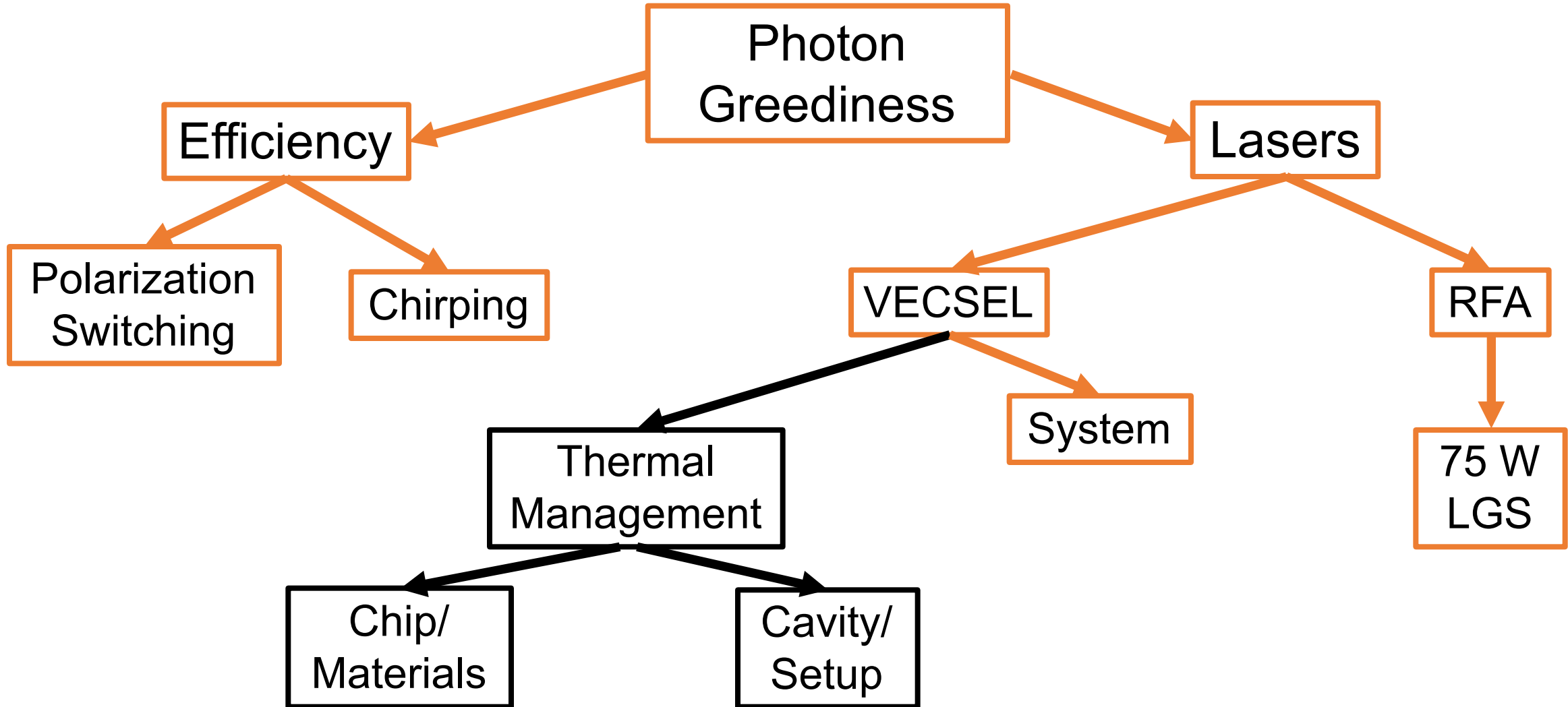


## 12W Multi-Mode Power at 589 nm

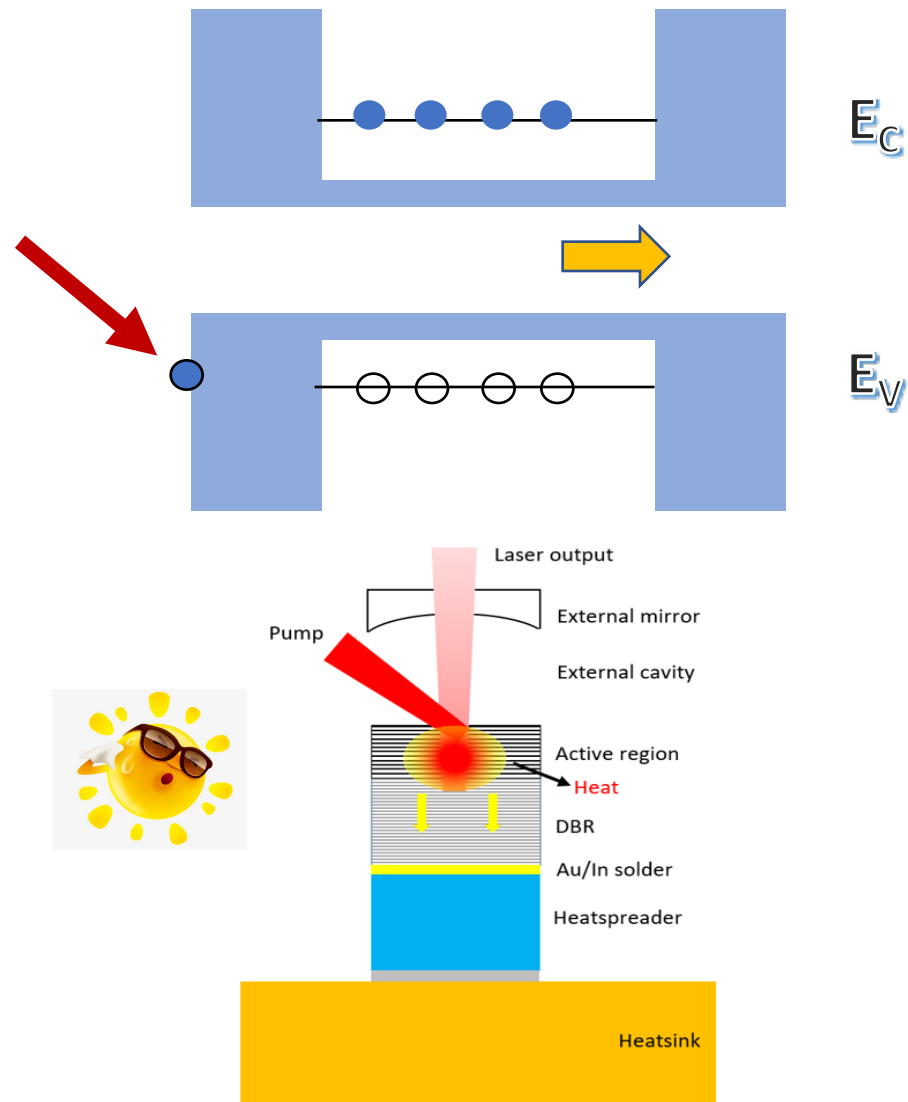


- Power emitted from two mirrors in the cavity – measurement at one end
- Higher single frequency is possible
  - Improved mode selectivity
  - Improved thermal management
  - Optimization of nonlinear crystal parameters

# Next Generation at SOR



# VECSEL technical overview



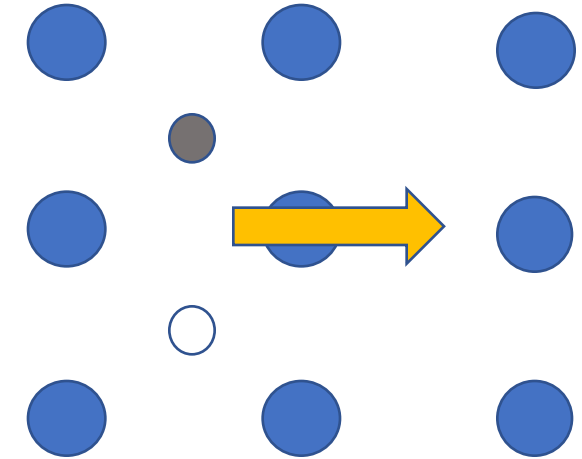
- Pump photons are absorbed causing generation of free electrons and holes in the quantum well barrier. Carriers diffuse to QW and recombine for photon emission.
- QW designed so discrete energy levels emit photons at wavelengths of interest
  - Bandgap energy of lasing materials
- Multiple Quantum Well (MQW)
- Amplification in gain region, usually using DBR
  - DBR-free adds additional external mirrors
  - Removes low thermal conductivity DBR
- Highly strained quantum well structures
  - Increase Indium, Increase Wavelength
  - InGaAs, large lattice constant mismatch, degradation
  - Addition of Phosphorus or Nitrogen

Fallahi, M., et al, IEEE Photonics Technol. Lett., 20, no. 20, 2008.  
 Kantola, E., et al, Optical Express, Vol 22. Issue 6, 6372-6380, March 2014.  
 Hackett, Shawn, UNM Digital Repository, November 2016

# Heat

Non-radiative recombination, growth defect, quantum defect

1. Non-radiative recombination: form of recombination in which a phonon is released
2. Growth defects: issues in epitaxial growth; MOCVD or MBE
3. Quantum defect: difference in the energy per photon between pump photons and emitted laser photons
4. Inefficient heat dissipation: DBR low thermal conductivity, soldering, SiC or diamond heatspreader



To chill or not to chill?

CONDENSATION

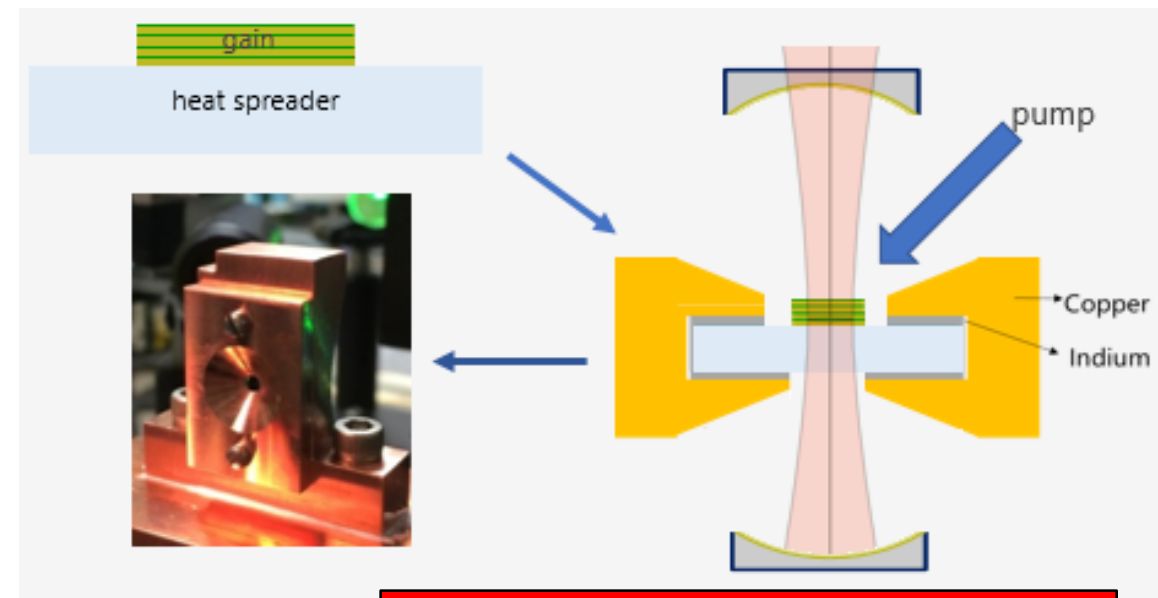
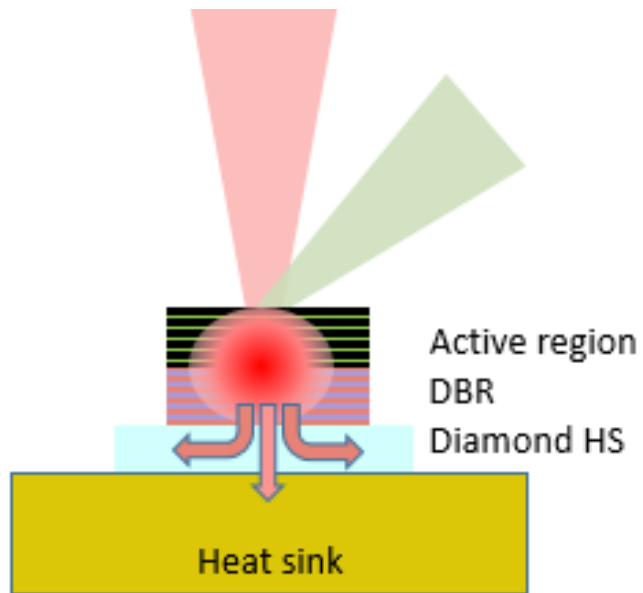
DIAMOND  
- Surface quality  
- Expensive for optical grade

SiC  
- Lower thermal conductivity

# Crystalline Mirror Solutions & University of New Mexico

## Elimination of the DBR in the VECSEL architecture

- Addition of a direct-bonded intracavity transparent heatspreader to MQW active region
- Comparing thermal conductivity and surface quality of diamond and SiC



Maximum 7 W reached at 1178nm

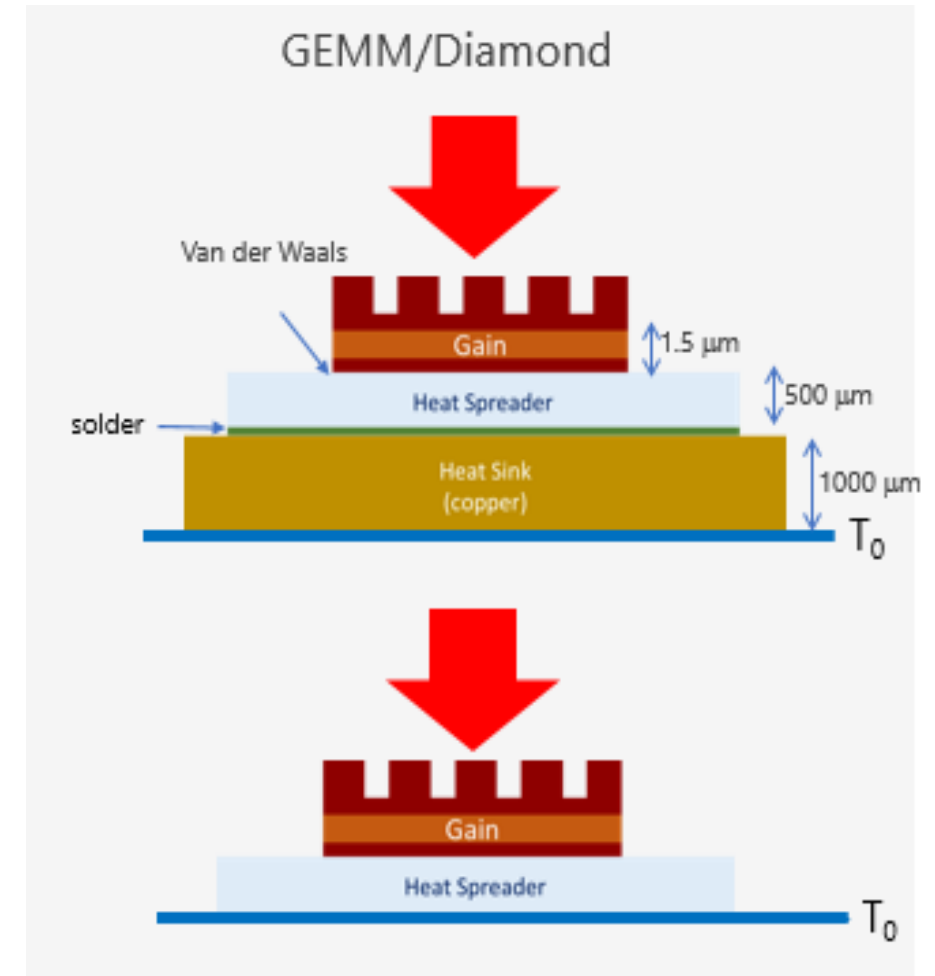
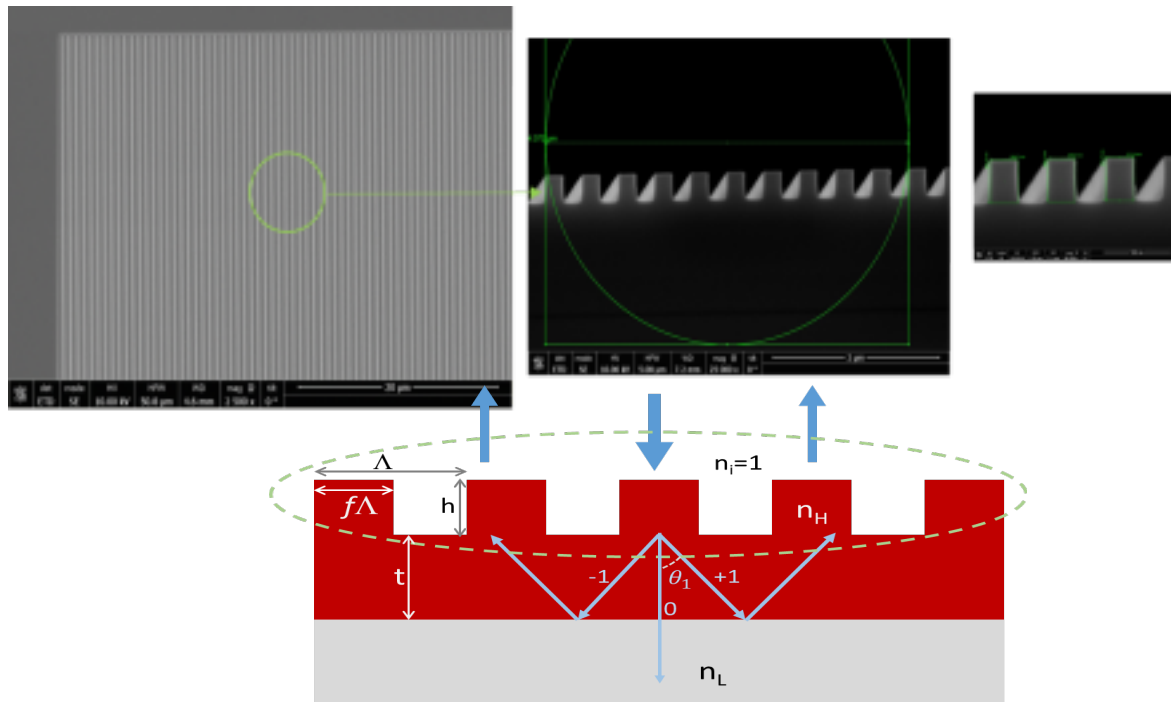
- M. Sheik-Bahae, U.S. Patent # 11/845,367 (2009)
- Z. Yang, A. R. Albrecht, J.G. Cederberg, M. Sheik-Bahae, *Opt. Express*, 23 (26), 33164 (2015)
- H. Kahle *et al.*, *Optica*, vol. 3, no. 12, pp. 1506–1512 (2016)
- S. Mirkhanov *et al.*, *Electron. Lett.*, vol. 53, no. 23, pp. 1537–1539, (2017)



# CMS & UNM cont.

Development of a novel VECSEL architecture:  
Gain-Embedded Meta Mirror technology (GEMM)

- Broadband active meta-mirror utilizes total internal reflection from a surface grating atop an underlying gain medium



- Z. Yang, D. Lidsky, M. Sheik-Bahae, "Gain-Embedded Meta Mirrors for Optically Pumped Semiconductor Disk Lasers" <https://arxiv.org/pdf/1901.00472> (2018)

# Big Thanks To...

- Sodium team:

Lee Kann, Mark Eickhoff, David Ireland

- All industry/research teams:

Areté and MIT/LL

HartScientific and University of Arizona

CMS and University of New Mexico



AFRL/RDS/Starfire Optical Range/RQ Fugate

Thank you!

QUESTIONS?